

Family behavior of the optical transition energies in single-wall carbon nanotubes of smaller diameters

Ge. G. Samsonidze

Department of Electrical Engineering and Computer Science, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139-4307

R. Saito, N. Kobayashi, A. Grüneis,^{a)} and J. Jiang

Department of Physics, Tohoku University and CREST JST, Aoba, Sendai 980-8578, Japan

A. Jorio

Departamento de Física, Universidade Federal de Minas Gerais, Belo Horizonte, MG 30123-970, Brazil

S. G. Chou

Department of Chemistry, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139-4307

G. Dresselhaus

Francis Bitter Magnet Laboratory, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139-4307

M. S. Dresselhaus^{b)}

Department of Electrical Engineering and Computer Science and Department of Physics, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139-4307

(Received 20 July 2004; accepted 5 October 2004)

Using the extended tight-binding model that allows bond lengths and angles to vary, the optical transition energies E_{ii} in single-wall carbon nanotubes are calculated as a function of inverse tube diameter. After geometrical structure optimization, the $2n+m=\text{constant}$ family behavior observed in photoluminescence (PL) experiments is obtained, and detailed agreement between the calculations and PL experiments is achieved after including many-body corrections. © 2004 American Institute of Physics. [DOI: 10.1063/1.1829160]

The electronic band structure of single-wall carbon nanotubes (SWNTs) is often obtained by applying the zone-folding scheme to the band structure of the graphene layer, where the latter is calculated within the tight-binding (TB) approximation.¹ The TB transfer and overlap integrals are then fitted to the results of experiments, such as resonance Raman scattering (RRS)² or scanning tunneling spectroscopy.³ While this approach provides reliable results for larger diameter SWNTs (>1.2 nm),⁴ it fails in the smaller diameter region (<1.2 nm), as has been shown recently in photoluminescence (PL) studies of SWNTs dispersed by a surfactant in an aqueous solution.⁵ An empirical fitting approach was developed by Weisman and Bachilo to reproduce the results of the PL experiments.⁶ In this letter, we present an extension of the zone-folding scheme and the TB approximation to the smaller diameter region which agrees well with the PL empirical fit,⁶ thus providing a theoretical basis to account for this empirical fit that can now be safely applied to many experiments. It can also be extended to predict electronic and optical properties of SWNTs over a wider SWNT diameter and energy range, as well as for metallic SWNTs that are missing from the PL empirical fit due to the quenching of the PL signal by metallic SWNTs.

The optical properties of SWNTs are determined by the electronic transitions between van Hove singularities (vHSs) in the density of states (DOS) arising from the one-dimensional (1D) structure of SWNTs. These transition en-

ergies E_{ii} between vHSs for SWNTs of different structural (n,m) indices are commonly summarized in the so-called Kataura plot that is widely used in RRS and PL studies of SWNTs.⁷ The Kataura plot depicts the E_{ii} as a function of SWNT diameters (d_t) or inverse SWNT diameters ($1/d_t$). For each (n,m) SWNT, $d_t = a\sqrt{n^2 + nm + m^2}/\pi$, where $a = \sqrt{3}a_{CC}$ is the graphene lattice constant, and $a_{CC} = 0.142$ nm is the C–C interatomic distance.¹ The E_{ii} energies in the Kataura plot are arranged in bands ($E_{11}^S, E_{22}^S, E_{11}^M$, etc.) for semiconducting (S) and metallic (M) SWNTs, respectively, where the index i enumerates the vHSs in the valence and conduction bands away from the Fermi level. Within each band in the Kataura plot, the E_{ii} energies observed from PL measurements follow “family” patterns for SWNTs with $2n+m=3p+r$, where p is an integer and $r=0,1,2$ define metallic, semiconducting type I (SI) and type II (SII) SWNTs, respectively. The PL empirical fit⁶ provides the first two optical transition energies for S SWNTs, E_{11}^S and E_{22}^S . When comparing the Kataura plot obtained from the PL empirical fit⁶ with the one calculated from the TB approximation,⁷ two major differences can be found. First, the experimental E_{22}^S/E_{11}^S ratio in the large d_t limit is less than 2, while the tight-binding E_{22}^S/E_{11}^S ratio approaches 2 with increasing d_t (the “ratio” problem). Second, the empirical spread of the E_{ii}^S energies within the same $2n+m=\text{constant}$ family is much larger than the corresponding spread of the TB E_{ii}^S energies at constant d_t (the “family spread” problem). While the “ratio” problem can be explained by many-body effects,⁸ the “family spread” is mainly attributed to the curvature effects and to the C–C bond length optimization in smaller d_t SWNTs,

^{a)}Present address: Institute for Solid State Research, Leibniz Institute for Solid State and Materials Research, 01171 Dresden, Germany.

^{b)}Electronic mail: millie@mgm.mit.edu

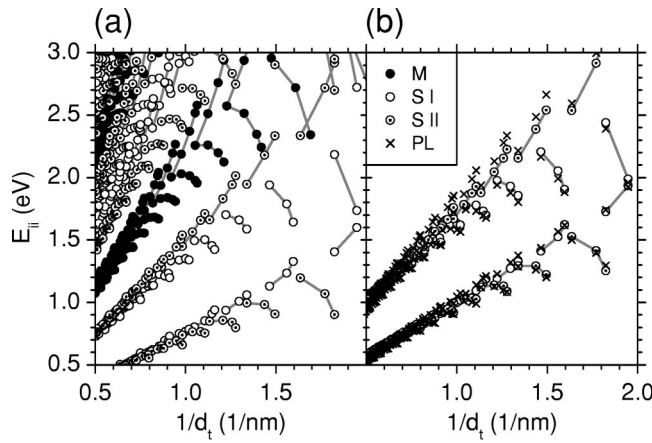


FIG. 1. (a) “Kataura” plot of transition energies E_{ii} vs inverse diameter $1/d_t$ for metallic (closed dots) and semiconducting type I (open dots) and type II (marked dots) SWNTs based on the extended tight-binding (ETB) model after geometrical structure optimization. (b) Comparison between the ETB calculations for E_{11}^S and E_{22}^S and the PL empirical fit (crosses)—Ref. 6—after making the many-body corrections of Eq. (1).

which are missing from the conventional TB approximation.¹ Meanwhile, it has been shown that long-range interactions of the π orbitals are not negligible,⁹ and curvature of SWNT sidewalls yields the sp^2 – sp^3 rehybridization in the small d_t limit. The curvature effect can be included in the TB model, as has been recently reported,¹⁰ by extending the basis set to the atomic s , p_x , p_y , and p_z orbitals that form the σ and π molecular orbitals according to the Slater–Koster formalism.¹ This extended tight-binding (ETB) model utilizes the TB transfer and overlap integrals as functions of the C–C interatomic distance calculated within a density-functional theory (DFT) framework,¹¹ thus including long-range interactions and bond-length variations within the SWNT sidewall. The atomic p -orbitals are aligned in the cylindrical coordinates of the SWNT sidewall according to the symmetry-adapted scheme¹⁰ (p_z is orthogonal to the SWNT sidewall, while p_x and p_y are parallel to the SWNT sidewall for each C atom), which allows us to consider an 8×8 Hamiltonian for the graphene unit cell of two C atoms (A and B), even for chiral SWNTs with large translational unit cells, thus greatly simplifying the calculations. Furthermore, the total energy of the SWNT can be calculated using the short-range repulsive potential obtained from DFT calculations,¹¹ and the geometrical structure optimization can then be performed.¹⁰ As shown elsewhere,¹⁰ the resulting optimized SWNT diameter d_t^{ETB} slightly exceeds its ideal value $d_t = a\sqrt{n^2 + nm + m^2}/\pi$. It is essential to utilize the optimized SWNT structure since the family spread in the Kataura plot is very sensitive to the relaxed atomic positions.

We have used the ETB model^{10,11} first to optimize the SWNT structure and then to calculate the vHSs in the DOS of SWNTs and finally to construct the Kataura plot which is shown in Fig. 1(a). By comparing it to the PL empirical fit,⁶ we found that the family spread observed in PL studies is closely reproduced by the ETB approximation. The ETB model thus provides the proper chirality dependence for the E_{ii} energies, since the SWNT chirality changes from armchair-like (A) to zigzag-like (Z) along $2n+m=\text{constant}$ family lines. The differences between the E_{ii}^{PL} energies obtained from PL measurements⁶ and the corresponding E_{ii}^{ETB} energies calculated from the ETB model, hereafter referred

to as ΔE_{ii} , thus weakly depend on the SWNT chirality (generally within the accuracy of PL measurements), while they show a monotonic dependence on d_t .

These energy differences can be explained by many-body effects, which consist of electron–electron Coulomb repulsion that upshifts E_{ii} and of the exciton binding that downshifts E_{ii} .¹² Because of the 1D SWNT structure, electron–electron Coulomb repulsion exceeds the exciton binding so that, overall, many-body E_{ii} energies are upshifted from one-electron E_{ii} energies.¹² Since the Coulomb interaction range in SWNTs is of the order of 10 nm,¹³ which is much larger than d_t , the many-body corrections to E_{ii} are weakly sensitive to the SWNT chirality but essentially only depend on d_t , on the subband index $i=1$ or 2 , and on the S type I or II. We thus fitted $\Delta E_{11}^{\text{SI}}$, $\Delta E_{11}^{\text{SII}}$, $\Delta E_{22}^{\text{SI}}$, and $\Delta E_{22}^{\text{SII}}$ as functions of d_t for all (n, m) SWNTs observed in PL studies.⁶ The fit yields

$$\Delta E_{11}^{\text{SI}} = E_{11}^{\text{SI PL}} - E_{11}^{\text{SI ETB}} = (0.15 + 0.11 \text{ nm}/d_t) \text{eV},$$

$$\Delta E_{11}^{\text{SII}} = E_{11}^{\text{SII PL}} - E_{11}^{\text{SII ETB}} = (0.12 + 0.11 \text{ nm}/d_t) \text{eV},$$

$$\Delta E_{22}^{\text{SI}} = E_{22}^{\text{SI PL}} - E_{22}^{\text{SI ETB}} = (0.25 - 0.03 \text{ nm}/d_t) \text{eV},$$

$$\Delta E_{22}^{\text{SII}} = E_{22}^{\text{SII PL}} - E_{22}^{\text{SII ETB}} = (0.31 - 0.03 \text{ nm}/d_t) \text{eV}. \quad (1)$$

We add ΔE_{11}^{S} and ΔE_{22}^{S} given by Eq. (1) to E_{11}^{S} and E_{22}^{S} calculated in the ETB model, and then plot the resulting E_{11}^{S} and E_{22}^{S} as a function of $1/d_t$ in Fig. 1(b). For comparison, E_{11}^{S} and E_{22}^{S} from the PL empirical fit⁶ are also plotted. One can see detailed agreement between the ETB calculations and the PL empirical fit once the many-body corrections of Eq. (1) are taken into account.

The families of $2n+m=\text{constant}$ bend downward with increasing $1/d_t$ in the smaller d_t region, as shown in Fig. 1(b), in full agreement with the PL empirical fit.⁶ When using the conventional π -band nearest-neighbor TB approximation, the $2n+m=\text{constant}$ families never bend down, but rather follow the same general tendency of a linear increase in E_{ii} as $1/d_t$ increases. It should be emphasized that when applying the ETB model for the nonoptimized SWNT structure, the calculated family spread is not nearly large enough to fit the spread observed experimentally. This indicates the importance of the geometrical structure optimization on the E_{ii} values.

To compare the SWNT structures optimized by using the ETB model with the results of other independent geometrical structure optimizations, we plot the change in the C–C bond lengths for each SWNT as a function of curvature ($1/d_t^2$) in Fig. 2. Similar calculations have been performed by Kanamitsu *et al.*¹⁴ for zigzag SWNTs using the DFT framework. Kanamitsu’s bond lengths are also shown for comparison with our calculations. We can see in Fig. 2 that the results of the two independent geometrical structure optimizations agree with each other and follow the same general pattern, yet some deviations are also present [for example, at $1/d_t^2 = 0.033 \text{ \AA}^{-2}$ for the (7,0) SWNT]. One of the two bond lengths for zigzag SWNTs increases with curvature while the other decreases, in agreement with the physical picture that the optimization process increases the SWNT diameter (from d_t to d_t^{ETB}) and shrinks the SWNT length. This optimization

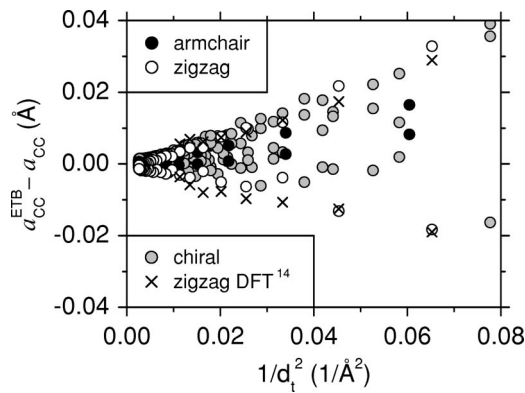


FIG. 2. Differences between the C–C bond lengths a_{CC}^{ETB} for each SWNT and $a_{CC}=0.142$ nm in the flat graphene layer as a function of curvature $1/d_t^2$. Open, closed, and gray dots denote the bond lengths of zigzag, armchair, and chiral SWNTs, respectively, calculated from the ETB model for the optimized SWNT structures. For comparison, crosses show the bond lengths of zigzag SWNTs from DFT calculations (Ref. 14).

becomes more important as d_t decreases.¹⁰ The SWNT structures optimized by using the ETB model are thus consistent with other available data.

Once the geometrical structure optimization is performed and the E_{ii} energies are calculated, we plot them as a function of $1/d_t$ in the Kataura plot as shown in Fig. 1. However, the Kataura plot used in RRS studies of SWNTs depicts the E_{ii} energies as a function of ω_{RBM} , the radial-breathing mode (RBM) Raman frequency, which is known to vary as $1/d_t$.¹⁵ The force-constants for the RBM and other phonon modes can also be calculated for the optimized SWNT structures using the ETB model. The total energy for each SWNT is first calculated and its second derivative is then taken for the atomic displacements along the phonon eigenvector (which corresponds to an increase in d_t in the case of the RBM) yielding the force-constant for this particular phonon mode. The calculated RBM frequencies ω_{RBM}^{ETB} are shown as a function of $1/d_t$ in Fig. 3. The calculated ω_{RBM}^{ETB}

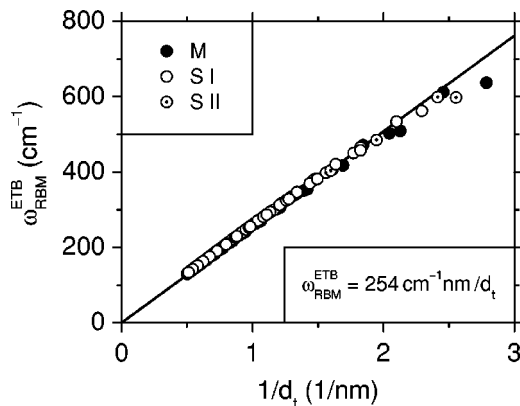


FIG. 3. The radial-breathing mode (RBM) frequencies ω_{RBM}^{ETB} for each SWNT as a function of inverse diameter $1/d_t$. The frequencies ω_{RBM}^{ETB} are calculated from the extended tight-binding (ETB) model for the optimized SWNT structures. Closed, open, and marked dots correspond to metallic, semiconducting type I, and type II SWNTs, respectively. The line shows a linear fit $\omega_{RBM}^{ETB}=254$ cm^{-1} nm/d_t to the calculated points.

follows a linear dependence with $1/d_t$ up to the smaller d_t region, where ω_{RBM}^{ETB} slightly downshifts from the linear behavior, in agreement with published results.¹⁵ By fitting the calculated ω_{RBM}^{ETB} , we obtain $\omega_{RBM}^{ETB}=254$ cm^{-1} nm/d_t . The proportionality coefficient 254 cm^{-1} nm is slightly higher than the values observed experimentally in RRS studies of SWNTs (223–248 cm^{-1} nm for different samples^{2,4,5}). The calculated force-constants are generally upshifted from experimental values according to the variational principle. Thus the ω_{RBM} calculated from first principles should generally be scaled down by $\sim 10\%$. Such scaling brings the calculated ω_{RBM}^{ETB} down to the experimental range of the observed ω_{RBM} . Using the proper RBM proportionality coefficient for a specific SWNT sample, the Kataura plot in Fig. 1 can be redrawn as a function of ω_{RBM} for its practical use in RRS studies. Furthermore, the geometrical structure optimization can be performed in the presence of different SWNT wrapping agents, thus allowing us to predict theoretically the observed changes in the E_{ii} energies for different SWNT samples.

In summary, a theoretical basis is provided in support of the PL empirical fit⁶ for the optical transition energies in SWNTs. The model may be extended with future experiments to a wider SWNT diameter range, to other electronic subbands beyond E_{11}^S and E_{22}^S , to metallic SWNTs not seen in PL studies, and to experimental SWNT samples containing a variety of surfactants and wrapping agents, and whether or not SWNTs are suspended, or are on particular substrates, and whether the SWNTs are solvated or not.

The MIT authors acknowledge support under NSF Grant No. DMR 04-05538. The Tohoku U. authors acknowledge a Grant-in-Aid (Nos. 13440091 and 16076201) from the Ministry of Education, Japan. A.J. acknowledges financial support from CNPq-Brazil (Profix).

¹R. Saito, G. Dresselhaus, and M. S. Dresselhaus, *Physical Properties of Carbon Nanotubes* (Imperial College Press, London, 1998).

²A. Jorio, R. Saito, J. H. Hafner, C. M. Lieber, M. Hunter, T. McClure, G. Dresselhaus, and M. S. Dresselhaus, *Phys. Rev. Lett.* **86**, 1118 (2001).

³M. Ouyang, J. L. Huan, C. L. Cheung, and C. M. Lieber, *Science* **292**, 702 (2001).

⁴A. G. Souza Filho, S. G. Chou, G. G. Samsonidze, G. Dresselhaus, M. S. Dresselhaus, L. An, J. Liu, A. K. Swan, M. S. Ünlü, B. B. Goldberg, A. Jorio, A. Grüneis, and R. Saito, *Phys. Rev. B* **69**, 115428 (2004).

⁵S. M. Bachilo, M. S. Strano, C. Kittrell, R. H. Hauge, R. E. Smalley, and R. B. Weisman, *Science* **298**, 2361 (2002).

⁶R. B. Weisman and S. M. Bachilo, *Nano Lett.* **3**, 1235 (2003).

⁷H. Kataura, Y. Kumazawa, Y. Maniwa, I. Umezu, S. Suzuki, Y. Ohtsuka, and Y. Achiba, *Synth. Met.* **103**, 2555 (1999).

⁸C. L. Kane and E. J. Mele, *Phys. Rev. Lett.* **90**, 207401 (2003).

⁹S. Reich, J. Maultzsch, C. Thomsen, and P. Ordejón, *Phys. Rev. B* **66**, 035412 (2002).

¹⁰V. N. Popov, *New J. Phys.* **6**, 17 (2004).

¹¹D. Porezag, T. Frauenheim, T. Köhler, G. Seifert, and R. Kaschner, *Phys. Rev. B* **51**, 12947 (1995).

¹²T. Ando, *J. Phys. Soc. Jpn.* **66**, 1066 (1997).

¹³C. D. Spataru, S. Ismail-Beigi, L. X. Benedict, and S. G. Louie, *Phys. Rev. Lett.* **92**, 077402 (2004).

¹⁴K. Kanamitsu and S. Saito, *J. Phys. Soc. Jpn.* **71**, 483 (2002).

¹⁵J. Kürti, V. Zólyomi, M. Kertesz, and G.-Y. Sun, *New J. Phys.* **5**, 125 (2003).